Improving noble gas based paleoclimate reconstruction and groundwater dating using $^{20}\text{Ne}/^{22}\text{Ne}$ ratios


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Abstract—The interpretation of noble gas concentrations in groundwater with respect to recharge temperature and fractionated excess gas leads to different results on paleo-climatic conditions and on residence times depending on the choice of the gas partitioning model. Two fractionation models for the gas excess are in use, one assuming partial re-equilibration of groundwater supersaturated by excess air (PR-model, Stute et al., 1995), the other assuming closed-system equilibration of groundwater with entrapped air (CE-model, Aeschbach-Hertig et al., 2000). In the example of the Continental Terminal aquifers in Niger, PR- and CE-model are both consistent with the data on elemental noble gas concentrations (Ne, Ar, Kr, and Xe). Only by including the isotope ratio $^{20}\text{Ne}/^{22}\text{Ne}$ it can be demonstrated that the PR-model has to be rejected and the CE-model should be applied to the data. In dating applications $^3\text{He}$ of atmospheric origin ($^3\text{He}_{atm}$) required to calculate $^3\text{H} - ^3\text{He}$ water ages is commonly estimated from the Ne excess presuming that gas excess is unfractiated air (UA-model). Including in addition to the Ne concentration the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio and the concentration of Ar enables a rigorous distinction between PR-, CE- and UA-model and a reliable determination of $^3\text{He}_{atm}$ and of $^3\text{H} - ^3\text{He}$ water ages. Copyright © 2002 Elsevier Science Ltd

1. INTRODUCTION

Noble gas concentrations in groundwater have been employed to provide information on paleo-temperatures (e.g., Aeschbach-Hertig et al., 2000; Andrews and Lee, 1979; Beyerle et al., 1998; Mazor, 1972; Stute et al., 1995; Weyhenmeyer et al., 2000) and on residence times in aquifers ($^3\text{H} - ^3\text{He}$ dating e.g., Schlosser et al., 1988; Schlosser et al., 1989; Tolstikhin and Kamenskiy, 1969; $^3\text{He}$ dating e.g., Andrews and Lee, 1979; Solomon, 2000). In addition to applications in paleo-climate and dating studies, noble gases can serve as excellent tools to study gas exchange between the atmosphere and groundwater because noble gases are not affected by biogeochemical transformations. The interpretation of noble gas concentrations in groundwater has been improved in the last years and rigorous least squares methods are now available (Aeschbach-Hertig et al., 1999; Ballentine and Hall, 1999) which allow the estimation of e.g., recharge temperature by considering the noble gases Ne, Ar, Kr, and Xe simultaneously.

In most aquifers noble gas concentrations are higher than the concentrations expected at atmospheric solubility equilibrium. This gas excess, commonly called “excess air” (Heaton and Vogel, 1981), must be considered in the calculation of recharge temperature (Stute and Schlosser, 1993) and in the calculation of tritiogenic $^3\text{He}$ for dating purposes (Schlosser et al., 1989). The name “excess air” however is misleading because in many aquifers (e.g., Aeschbach-Hertig et al., 2000; Stute et al., 1995) the gas excess does not have the same composition as atmospheric air but the heavier noble gases are enriched compared to the lighter ones. In several aquifers the amount and composition of the gas excess varies systematically over time and might provide additional information on past climatic conditions related to changes in recharge dynamics (Aeschbach-Hertig et al., 2001; Stute and Talma, 1998, Beyerle et al., in press).

Although gas excess appears to be very common in groundwaters its formation is still not well understood in detail. Two different simplified models describing the formation of excess air and its fractionation (Aeschbach-Hertig et al., 2000; Stute et al., 1995) have been employed to account for the excess of noble gases in paleo-climatic studies and in $^3\text{H} - ^3\text{He}$ dating applications. The two gas exchange models differ in their assumptions on how excess air is formed and on the mechanisms responsible for fractionation.

Because the choice of the gas exchange model affects the results on paleo-temperatures and on $^3\text{H} - ^3\text{He}$ water ages, it is necessary to decide on the basis of the available data which of the two models should be applied. However, in most groundwaters the number of elemental noble gas concentrations which can be employed in the inverse fitting procedure to simultaneously estimate recharge temperature, gas excess and its fractionation is limited to four, namely Ne, Ar, Kr and Xe. Unfortunately He, which is the most sensitive of all noble gases to the choice of the gas exchange model, usually cannot be included in the fitting procedure because it has significant additional non-atmospheric sources which cannot be quantified independently.

In this study we demonstrate that the empirical constraints on the gas exchange models can be improved significantly by considering the concentrations of noble gas isotopes and isotope ratios in addition to the commonly used elemental con-
Table 1a. Partial pressure in the atmosphere, equilibrium concentration $C^*$ and molecular diffusivities $D$ of noble gases and their isotopes at different temperatures. Equilibrium concentrations are given for freshwater with $S = 0$ at $P = 1$ atm of moist air. $C^*$ are calculated from the solubilities for $^4$He, $^3$He, Ne, Ar, and Kr (Benson and Krause, 1980; Weiss, 1970; Weiss and Kyser, 1978), and the solubilities for Xe (Cleaver, 1979) using the correction for moist air (Aeschbach-Hertig et al., 1999). $C^*$ of $^{22}$Ne and $^{40}$Ar are calculated from the equilibrium concentration of Ne and Ar using the fractionation factors given by Beyerle et al. (2000a). Solubilities of $^{20}$Ne and $^{40}$Ar are assumed to correspond to the solubilities of Ne and Ar, respectively. Volume fractions in dry air are taken from Oshima and Podosek (1983) and molecular diffusivities were calculated from the empirical relations of Jähne et al. (1987). Diffusivities of Ar, $^{36}$Ar, $^{32}$Ne and $^3$He were calculated from the other noble gas diffusivities assuming that the $D$ is inversely proportional to the square root of the atomic mass.

<table>
<thead>
<tr>
<th>$C^*$ at 5°C</th>
<th>$C^*$ at 10°C</th>
<th>$C^*$ at 20°C</th>
<th>$C^*$ at 30°C</th>
<th>$D$ at 5°C</th>
<th>$D$ at 10°C</th>
<th>$D$ at 20°C</th>
<th>$D$ at 30°C</th>
</tr>
</thead>
<tbody>
<tr>
<td>(cm$^3$ STP g$^{-1}$)</td>
<td>(10$^{-9}$ m$^2$s$^{-1}$)</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>He, $^4$He</td>
<td>5.24 · 10$^{-6}$</td>
<td>4.760 · 10$^{-8}$</td>
<td>4.645 · 10$^{-8}$</td>
<td>4.477 · 10$^{-8}$</td>
<td>4.357 · 10$^{-8}$</td>
<td>5.195</td>
<td>5.680</td>
</tr>
<tr>
<td>Ne</td>
<td>1.818 · 10$^{-3}$</td>
<td>2.123 · 10$^{-7}$</td>
<td>2.017 · 10$^{-7}$</td>
<td>1.851 · 10$^{-7}$</td>
<td>1.724 · 10$^{-7}$</td>
<td>2.627</td>
<td>2.942</td>
</tr>
<tr>
<td>Ar</td>
<td>9.34 · 10$^{-3}$</td>
<td>4.359 · 10$^{-4}$</td>
<td>3.861 · 10$^{-4}$</td>
<td>3.119 · 10$^{-4}$</td>
<td>2.599 · 10$^{-4}$</td>
<td>1.646</td>
<td>1.870</td>
</tr>
<tr>
<td>Kr</td>
<td>1.14 · 10$^{-3}$</td>
<td>1.057 · 10$^{-7}$</td>
<td>9.103 · 10$^{-8}$</td>
<td>6.967 · 10$^{-8}$</td>
<td>5.523 · 10$^{-8}$</td>
<td>1.029</td>
<td>1.200</td>
</tr>
<tr>
<td>Xe</td>
<td>8.7 · 10$^{-3}$</td>
<td>1.581 · 10$^{-8}$</td>
<td>1.318 · 10$^{-8}$</td>
<td>9.517 · 10$^{-9}$</td>
<td>7.178 · 10$^{-9}$</td>
<td>0.788</td>
<td>0.929</td>
</tr>
<tr>
<td>$^{20}$Ne</td>
<td>1.645 · 10$^{-5}$</td>
<td>1.921 · 10$^{-7}$</td>
<td>1.826 · 10$^{-7}$</td>
<td>1.675 · 10$^{-7}$</td>
<td>1.560 · 10$^{-7}$</td>
<td>2.627</td>
<td>2.942</td>
</tr>
<tr>
<td>$^{40}$Ar</td>
<td>9.303 · 10$^{-3}$</td>
<td>4.342 · 10$^{-4}$</td>
<td>3.846 · 10$^{-4}$</td>
<td>3.107 · 10$^{-4}$</td>
<td>2.589 · 10$^{-4}$</td>
<td>1.646</td>
<td>1.870</td>
</tr>
</tbody>
</table>

Concentrations of noble gases. In particular, including Ne isotopes proves to be very useful in distinguishing between different gas exchange models. The interpretation of data on noble gas isotopes with respect to excess air and its fractionation is demonstrated using noble gas data from the Continental Terminal aquifers in Niger (Beyerle et al., 2002).

2. THEORY AND METHODS

2.1. Gas Exchange Models and their Sensitivity to Noble Gas Concentrations

The most simple explanation for the excess of dissolved atmospheric gases in groundwater is that initially the gas concentrations in the groundwater are in equilibrium with the atmosphere and that gas bubbles with atmospheric gas composition are trapped and completely dissolved, introducing excess gases in the same ratio as in the atmosphere into the water. This assumption, which we refer to as the unfractionated excess air (UA) model, has been used in most paleoclimatic studies based on noble gas concentrations from groundwaters (e.g., Andrews and Lee, 1979; Heaton and Vogel, 1981; Stute and Schlitzer, 1993). However, recent studies have shown that the gas excess can be fractionated compared to the atmospheric gas composition. Two models have been developed to describe this situation, the partial re-equilibration model (PR-model) by Stute et al. (1995), and the closed system equilibration model (CE-model) by Aeschbach-Hertig et al. (2000). In the following we briefly describe the underlying model assumptions and their consequences on the noble gas concentrations and fractionation.

The PR-model assumes that initially the trapped air bubbles dissolve completely, as in the UA-model. The initial excess air suggested by the model can correspond to gas concentrations which are several times larger than the atmospheric equilibrium concentrations (e.g., Stute et al., 1995). This also implies that, initial oxygen levels in groundwater might be significantly above atmospheric equilibrium which would have consequences on groundwater ecology and water quality. According to the model, the gas excess in the water leads to a diffusion of gas out of the water across the groundwater table and diffusion is assumed to be at molecular level. Because the molecular diffusivities of the noble gases decrease with atomic mass, the light noble gases are lost much faster than heavier ones (Table 1a). This process increases the ratio of heavy to light noble gas concentrations in the water and thus leads to the typical fractionation pattern of the noble gas composition in the gas excess of groundwater. Some of the fractionated gas excess remains in the groundwater because infiltration and groundwater flow eventually move the water away from the air/water interface, preventing further gas loss by diffusion.

Conceptually, the model assumes a partial re-equilibration of water which was initially over-saturated by air with atmospheric composition. The effect of the PR-model on the dissolved gas concentration $i$ can be described by (Aeschbach-Hertig et al., 1999, reformulation of the model by Stute et al., 1995):

$$C_i(T, S, P, A_{pr}, F_{pr}) = C^*_i(T, S, P) + A_{pr} \cdot z_i \cdot \exp\left(-F_{pr} \frac{D_i}{D_{Ne}}\right)$$

where $C^*_i$ is the equilibrium concentration between atmosphere and water at atmospheric pressure $P$, water temperature $T$ and salinity $S$ during air water partitioning. The equilibrium concentration is determined using the Henry coefficient of the gas considered. Details on the calculation of $C^*_i$ are given in Aeschbach-Hertig et al. (1999). $z_i$ is the atmospheric volume fraction of noble gas $i$ in dry air, $D_i$ is its molecular diffusivity, and $D_{Ne}$ is the molecular diffusivity of Ne. $A_{pr}$ is the amount of initial excess air and $F_{pr}$ is the fractionation parameter which must be positive. $F_{pr}$ can be interpreted as being proportional to the time during which the initial gas excess is lost to the atmosphere. In most cases $T$, $A_{pr}$ and $F_{pr}$ are treated as fit parameters whereas $P$ and $S$ are commonly prescribed as present day atmospheric pressure and present day salinity of the infiltrating water (for meteoric water $S=0$). Note, that for $F_{pr} = 0$ (1) reduces to the UA-model.

As in the UA- and PR-models, Aeschbach-Hertig et al. (2000) assume that the gas-concentrations in the recharging water initially are in equilibrium with the atmosphere at atmospheric pressure and soil temperature and that bubbles of at-
Table 1b. Ratios of partial pressure in the atmosphere, of equilibrium concentration and of molecular diffusivity at different temperatures for selected noble gas isotopes and Ar and Ne.

<table>
<thead>
<tr>
<th>Gas Ratio</th>
<th>5°C (–)</th>
<th>10°C (–)</th>
<th>20°C (–)</th>
<th>30°C (–)</th>
<th>D2/D8 at 5°C (–)</th>
<th>D2/D8 at 10°C (–)</th>
<th>D2/D8 at 20°C (–)</th>
<th>D2/D8 at 30°C (–)</th>
</tr>
</thead>
<tbody>
<tr>
<td>3He/4He</td>
<td>1.384 · 10^{-6}</td>
<td>1.3596 · 10^{-6}</td>
<td>1.3602 · 10^{-6}</td>
<td>1.3614 · 10^{-6}</td>
<td>1.3625 · 10^{-6}</td>
<td>1.155</td>
<td>1.155</td>
<td>1.155</td>
</tr>
<tr>
<td>36Ar/36Ar</td>
<td>3.380 · 10^{-3}</td>
<td>3.380 · 10^{-3}</td>
<td>3.380 · 10^{-3}</td>
<td>3.380 · 10^{-3}</td>
<td>3.380 · 10^{-3}</td>
<td>1.054</td>
<td>1.054</td>
<td>1.054</td>
</tr>
<tr>
<td>Ne/Ar</td>
<td>1.946 · 10^{-3}</td>
<td>4.870 · 10^{-4}</td>
<td>5.225 · 10^{-4}</td>
<td>5.935 · 10^{-4}</td>
<td>6.633 · 10^{-4}</td>
<td>1.596</td>
<td>1.573</td>
<td>1.532</td>
</tr>
</tbody>
</table>

Note: Table 1b represents the ratios of partial pressure in the atmosphere, equilibrium concentrations, and molecular diffusivities for selected noble gas isotopes at different temperatures. The table includes ratios for 3He/4He, 20Ne/22Ne, and Ne/Ar, with values presented at 5°C, 10°C, 20°C, and 30°C. The table also includes the molecular diffusivity ratios at 5°C, 10°C, 20°C, and 30°C. The table is used to improve paleoclimate reconstruction and groundwater dating.
Fig. 1. $^{20}\text{Ne}/^{22}\text{Ne}$ ratio as function of $A_{pr}$ and $F_{pr}$ for the PR-model (a) and $A_{ce}$ and $F_{ce}$ for the CE-model (b). Each line represents the variation of $^{20}\text{Ne}/^{22}\text{Ne}$ ratio with $F_{pr}$ (a) and $F_{ce}$ (b) at a fixed value of $A$, i.e., a fixed $A_{pr}$ and $A_{ce}$, respectively. Conditions at recharge are assumed to be $T = 30^\circ\text{C}$, $p = 1\text{atm}$ and $S = 0$. Note that the CE-model only allows $F_{ce}$ to range between 0 and 1.
ployed. The presence of terrigenic and radiogenic/tritiogenic He sources prevents the use of He to distinguish between the PR- and the CE-model.

Isotopes of the same noble gas have very similar Henry coefficients but can have significantly different molecular diffusivities (Table 1b). That is especially the case for the light noble gas isotopes 3He - 4He, 20Ne - 22Ne, and 36Ar - 40Ar. According to the above discussion, the isotopic ratios 3He/4He, 20Ne/22Ne and 36Ar/40Ar in the gas excess should remain between the atmospheric ratio and the ratio at atmospheric solubility equilibrium if fractionation occurs according to the CE-model, but might be significantly smaller if the fractionation depends on the differences in molecular diffusivities (PR-model). Because the 3He/4He ratio is affected by radiogenic/tritiogenic and terrigenic sources and the 36Ar/40Ar ratio is also influenced by terrigenic sources in aquifers with very large groundwater ages (e.g., Beyerle et al., 2000b; Torgersen et al., 1989), the 20Ne/22Ne ratio is the most reliable noble gas isotope ratio to test the fractionation models. The 36Ar/38Ar ratio could in principle also be used, although it is less sensitive to fractionation than the 20Ne/22Ne ratio, but to our knowledge it has never been measured in groundwater studies, due to the high experimental requirements (good mass resolution needed to separate 38Ar from 40Ar). Although molecular diffusivities differ much more between heavy noble gases and Ne than between 20Ne and 22Ne (Table 1), the 20Ne/22Ne ratio is best suited to distinguish between the PR- and the CE-model because its dependence on recharge temperature is negligible and the 20Ne/22Ne ratio can be measured with a much higher precision (typically 0.2%) than the concentrations of elemental noble gases (typically 1 to 2%).

2.2. Analysis of Noble Gas Data using the Software NOBLE

Noble gas data analysis with respect to recharge temperature, excess air and its fractionation is commonly performed by inverse fitting based on χ²-minimization, which takes the error of the measurements into account (Aeschbach-Hertig et al., 1999). The new fitting program Noble developed for this study extends the original fitting procedure of Aeschbach-Hertig et al. (1999) by incorporating the CE-model and including the possibility to consider not only elemental noble gas concentrations but also concentrations of noble gas isotopes and noble gas isotope ratios. Commonly the three unknown parameters T, A_p, F_p or T, A_c, F_c, respectively, are estimated from 4 observables, i.e., the concentration of Ne, Ar, Kr and Xe. Pressure P and salinity S at recharge are prescribed. Considering the 20Ne/22Ne and/or the 36Ar/40Ar ratios in addition to the elemental concentrations increases the number of observables and hence allows to estimate additional parameters (e.g., P if the location of recharge is unknown) or can be used to improve the constraint on the fit parameters.

Noble extends the original parameter set employed in Aeschbach-Hertig et al. (1999) by including optional parameters that describe the concentrations of terrigenic 3He and tritiogenic 4He, as well as the terrigenic 3He/4He ratio. An additional scaling parameter enables analysis based on the relative concentrations of the noble gases if the scaling of the absolute concentrations is unknown (e.g., unknown sample weight). A similar scaling technique has been used to interpret noble gas data from natural gas deposits where the degassing water volume is unknown (Ballentine et al., 1999). Furthermore, Noble supports ensemble fitting, i.e., the model parameters can be fitted not only to the gas concentrations of individual water samples separately but also to the concentrations from an ensemble of samples simultaneously. Therefore some of the fit parameters can be assumed to have the same value for all samples while other parameters may differ between samples. Propagation of the error of the measurements is performed by linear error propagation and optionally by a Monte-Carlo procedure to provide errors of the fit parameters. Both methods assume that measurement errors are normally distributed and independent. Noble calculates the probability p(χ²) for the minimum χ² obtained from the fitting to be observed at the given number of freedoms. This provides the possibility to apply a χ²-test to the models. Noble provides the probability p(χ²) for each water sample separately and for the entire data set p(χ² corr), where χ² corr is the sum of the χ² values obtained for the samples of the data set. NOBLE will be made available on the Internet (http://www.eawag.ch/research_e/w+s+t/UI/noblegasmethod.html).

3. STUDY AREA AND DATA

Details on the study area and an extended data set are given in Beyerle et al. (2002). In brief, groundwater samples were taken from the Continental Terminal (CT) aquifers located in south-western Niger, Africa, between 12.5 to 14.5° N and 2.5 to 4.5° E (Fig. 2). Within the CT formation three different aquifer compartments can be distinguished (CT3, CT2, CT1). CT3 is mostly unconfined, whereas CT2 and CT1 are confined. The Continental Terminal aquifer system is separated against the underlying Continental Intercalaire aquifer (CI) by massive low-permeable sediments mainly consisting of clays (Andrews et al., 1994). The altitude of the investigated area is almost constant increasing only slightly from 200 m to 300 m a.s.l. in north-easterly direction. The recharge area of the CT2 aquifer is located in the northern part of the study area and has an average altitude of 300 m a.s.l whereas the recharge areas of the CT1 and CI aquifers are further to the north at a mean altitude of ~350 to 400 m.

Groundwater samples were collected in April 1999 and June 2000 from drinking water wells either operated by manual pumps or by electric submersible pumps. In all samples, concentrations of the noble gas isotopes 3He, 4He, 20Ne, 22Ne, 36Ar, 40Ar, 86Kr and 136Xe were analysed according to the methods described in Beyerle et al. (2000a).

4. RESULTS AND DISCUSSION

4.1. Comparison of Gas Exchange Models

Figure 3 compares the model results from the PR- and the CE-model on recharge temperature for the CT aquifers derived by inverse fitting of the concentrations of Ne, Ar, Kr and Xe. Recharge temperatures are displayed as function of terrigenic 3He/4He, which can be interpreted as a qualitative measure of groundwater residence time because 3He/4He (here of crustal origin) accumulates over time. 3He/4He is the difference between the measured 3He concentration and the concentration of 3He of atmospheric origin (3He atm) determined from the PR- and the CE-model, respectively. The concentration of atmospheric
origin is defined as the sum of two contributions: the atmospheric equilibrium concentration at $T$, $S$ and $P$ and the concentration due to fractionated excess air.

The model assuming unfractionated excess air has to be rejected according to the $\chi^2$-test, making it necessary to account for fractionation. Both the PR-model and the CE-model predict Ne, Ar, Kr and Xe concentrations compatible with the data. The PR-model (open symbols, Fig. 3) yields systematically higher recharge temperatures than the CE-model (solid symbols, Fig. 3), on average by $1.5^\circ$C and for the 3 oldest samples even by $4^\circ$C. The difference in recharge temperature between young samples ($^{40}$He$_{ter}<1.5 \times 10^{-7}$ cm$^3$STPg$^{-1}$; $n=17$) and intermediate samples ($1.5 \times 10^{-7}$ cm$^3$STPg$^{-1}<^{40}$He$_{ter}<1.5 \times 10^{-6}$ cm$^3$STPg$^{-1}$; $n=7$) is about the same for the PR-model ($\Delta T = 4.0^\circ$C $\pm 1.8^\circ$C) and the CE-model ($4.8^\circ$C $\pm 1.3^\circ$C). A thorough discussion of paleo-climatic conditions in Niger based on a more comprehensive data set on tracers from the CT Aquifers is given elsewhere (Beyerle et al., 2002). Here we focus on the use of Ne isotopes to distinguish between the two models.

Figure 4 compares the measured $^{20}$Ne/$^{22}$Ne (Fig. 4a) and $^{36}$Ar/$^{40}$Ar ratios (Fig. 4b) with those predicted from the PR-model (open symbols) and the CE-model (solid symbols). The model parameters were derived by inverse fitting of Ne, Ar, Kr and Xe concentrations. The agreement between data and model prediction is significantly better for the CE-model than for the PR-model. Within error almost all measured $^{20}$Ne/$^{22}$Ne ratios ($^{20}$Ne/$^{22}$Ne$_{m}$) fall in the range between the isotopic ratio in the atmosphere (Ozima and Podosek, 1983) (Fig. 4a dashed lines, $z_{20Ne}/z_{22Ne}$) and the isotopic ratio at atmospheric equilibrium (Beyerle et al., 2000a) (Fig. 4a dotted lines, $^{20}$Ne$^*/^{22}$Ne$^*$). Considering the standard deviation of all data, the mean of the measured $^{20}$Ne/$^{22}$Ne and $^{36}$Ar/$^{40}$Ar ratios are consistent.
with a value between the corresponding ratios in the atmosphere and at atmospheric equilibrium (Table 2).

As discussed above, fractionation described by the CE-model always predicts isotopic ratios in the range between the values for air and atmospheric equilibrium. The PR-model however may predict isotopic ratios below this range as is demonstrated in the case of the CT Aquifers (Fig. 4 open symbols, Table 2).

The constraint on the fit-parameters $T$, $A$ and $F$ can be increased if the measured $^{20}\text{Ne}/^{22}\text{Ne}$ ratio is included as an observable in addition to the noble gas concentrations of Ne, Ar, Kr, and Xe. Using this extended data set, the CE-model yields a good fit ($p(\chi^2_{\text{CE}}=67) \approx 11\%$), whereas the PR-model has to be rejected ($p(\chi^2_{\text{PR}}=440) << 1\%$). This result reflects the fact that the measured $^{20}\text{Ne}/^{22}\text{Ne}$ ratios agree better with the predictions according to the CE-model than with the predictions according to the PR-model.

The best fit values for the parameters of the CE-model derived for the extended data set with the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio included are essentially the same as those derived from the elemental concentrations of Ne, Ar, Kr, and Xe only. In contrast, including the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio reduces the best fit values for $F_{\mu}$ of the PR-model by a factor of 10, resulting in $F_{\mu}$ ranging from 0 to 0.27 with a mean of $F_{\mu} = 0.1 \pm 0.1$. This implies that the predicted fractionation is small and that the PR-model essentially reduces to the UA-model. The constraint provided by the measured Ne isotopes ratios rules out significant diffusive gas loss.

Because the PR-model is incompatible with the data when considering measured $^{20}\text{Ne}/^{22}\text{Ne}$ ratios, only the CE-model adequately describes the gas exchange in the CT aquifers. Thus in case of the CT aquifers the parameters describing recharge temperature, gas excess and its fractionation derived by inverse fitting of Ne, Ar, Kr, Xe and the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio using the CE-model should be employed in the discussion of paleoclimatic change and for the calculation of $^3\text{He}_{\text{atm}}$ and $^4\text{He}_{\text{atm}}$ for dating purposes.

Observed $^{20}\text{Ne}/^{22}\text{Ne}$ ratios fall within the range between the ratio in the atmosphere and the ratio at atmospheric equilibrium not only in the CT aquifers of Niger but also in other aquifers where fractionated excess gas has been reported. Aeschbach-Hertig et al. (2000) analysed four data sets, but Ne isotope data are unfortunately available only for two of them (Belgium and Oman). In both cases no significant fractionation of the Ne isotopes was observed, e.g., in northern Oman (Weyhenmeyer et al., 2000) the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio measured in 9 samples ranges between 9.763 and 9.865 with a mean of 9.795 $\pm$ 0.029. This supports the applicability of the CE-model in noble gas based paleo-studies and for groundwater age dating. However, it does not necessarily exclude the PR-model. The PR- and the CE-model only describe the mean effect of gas exchange on noble gases and do not resolve details on the interaction between gas excess, entrapped air and recharging groundwater. Therefore, it is still unclear which environmental conditions lead to the noble gas pattern described by the PR-model and which result in the noble gas pattern described by the CE-model. Thus,
Fig. 4. Comparison of measured and predicted $^{20}\text{Ne}/^{22}\text{Ne}$ and $^{36}\text{Ar}/^{40}\text{Ar}$ ratios. Predictions were calculated from the model parameters describing recharge temperature, excess air and its fractionation derived with the CE-model (solid symbols) and PR-model (open symbols), respectively. Model parameters were obtained by inverse fitting to the concentrations of Ne, Ar, Kr and Xe.
Table 2. Mean measured and predicted $^{20}\text{Ne}^{32}\text{Ne}$ and $^{36}\text{Ar}^{40}\text{Ar}$ ratios in comparison with the corresponding ratios in the atmosphere and in atmospheric equilibrium.

\[
\begin{array}{|c|c|c|}
\hline
\text{Ratio} & \text{Mean of measurements} & \text{Mean of predicted ratios using the} \\
& (9.786 \pm 0.056) & \text{PR-model} \\
& (3.380 \pm 0.068) & \text{CE-model} \\
\hline
\hline
\end{array}
\]

whether the PR- or the CE-model should be applied in a specific aquifer can be judged only from data on concentrations of noble gases and their isotopes.

4.2. Dating Applications

Groundwater dating with the $^3\text{H}^\text{3He}$ method requires estimation of tritogenic $^3\text{He}$ ($^3\text{He}_{\text{atm}}$) by subtracting $^3\text{He}_{\text{atm}}$ and terrigenic $^3\text{He}$ ($^3\text{He}_{\text{ter}}$) from the measured $^3\text{He}$ ($^3\text{He}_{\text{e}}$) concentrations. Assuming that $^4\text{He}_{\text{atm}}$, $^4\text{He}_{\text{atm}}$ and the $^3\text{He}^\text{4He}$ ratio of the terrigenic component ($^4\text{He}^\text{4He}_{\text{ter}}$) is known, $^3\text{He}_{\text{atm}}$ can be calculated from the balance of $^4\text{He}$ and $^3\text{He}$ (e.g., Schlosser et al., 1989):

\[
^4\text{He}_{\text{atm}} = ^4\text{He}_{\text{e}} - ^4\text{He}_{\text{atm}} - ^3\text{He}_{\text{atm}} - ^3\text{He}_{\text{ter}} \cdot \left(\frac{^3\text{He}^\text{4He}_{\text{ter}}}{^3\text{He}^\text{4He}_{\text{atm}}}\right)
\]

4.2.1. Estimation of $^3\text{He}_{\text{atm}}$ and $^4\text{He}_{\text{atm}}$

The $^3\text{He}$ and $^4\text{He}$ components of atmospheric origin can be estimated from the $^3\text{He}$ and $^4\text{He}$ concentrations in the atmosphere (Ozima and Podosek, 1983) together with the $^4\text{He}$ and $^3\text{He}$ solubilities (Benson and Krause, 1980; Weiss, 1971) by employing an appropriate gas exchange model (e.g., the PR- or the CE-model) where model parameters describing gas excess and fractionation have been estimated by inverse fitting of noble gas concentrations. In case of the CT aquifers in Niger estimated $^3\text{He}_{\text{atm}}$ and $^4\text{He}_{\text{atm}}$ differ substantially between the PR- and CE-model if the inverse fitting procedure is based only on the elemental concentrations of Ne, Ar, Kr and Xe (Fig. 5a, Table 3).

If the measured $^{20}\text{Ne}^{32}\text{Ne}$ ratio is included as additional observable, $^3\text{He}_{\text{atm}}$ and $^4\text{He}_{\text{atm}}$ predicted with the PR-model increase substantially and closely agree with predictions assuming unfractonated excess air (Table 3). Predictions of $^3\text{He}_{\text{atm}}$ and $^4\text{He}_{\text{atm}}$ based on the CE-model are independent of whether the $^{20}\text{Ne}^{32}\text{Ne}_{\text{atm}}$ is included in the parameter estimation or not (see Table 3). Because only the CE-model is compatible with the measured elemental noble gas concentrations and the $^{20}\text{Ne}^{32}\text{Ne}$ ratio (see above) the $^3\text{He}_{\text{atm}}$ and $^4\text{He}_{\text{atm}}$ predicted with the CE-model are assumed to correspond to the true values.

4.2.2. Estimation of ($^4\text{He}/^3\text{He})_{\text{ter}}$ and natural (pre-bomb) Tritium

By restriction to samples that are not influenced by bomb tritium, the values of ($^4\text{He}/^3\text{He})_{\text{ter}}$ and the $^3\text{He}_{\text{ter}}$ of prebomb origin can be estimated by inverse fitting utilising the ensemble fitting option of Noble. To this end, noble gas concentrations and isotope ratios from all intermediate and old groundwater samples ($^4\text{He}_{\text{ter}} > 1.5 \times 10^{-7}$ cm$^3$STPg$^{-1}$; 10 samples) were compiled into a data ensemble. The fit-parameters $T$, $A$ and $F$ determining the noble gas concentration of atmospheric origin and the fit parameter describing the concentration of $^3\text{He}_{\text{ter}}$ are assumed to be different for each water sample. ($^4\text{He}/^3\text{He})_{\text{ter}}$ and prebomb $^3\text{He}_{\text{ter}}$ are assumed to be the same for all water samples, i.e., for the data ensemble.
For this data set the CE-model gives \((\HeHe)_{\text{atm}} = (2.2 \pm 0.1) \times 10^{-8}\) and \(\HeHe = 0.2 \pm 0.3\) TU (1 cm\(^3\)STPg\(^{-1}\) = 4.019 \times 10^{14}\) TU). The value of \(\HeHe\) corresponds to the concentration of prebomb tritium at the groundwater table. Prebomb \(^3\)H close to zero is reasonable because radioactive decay during the water transport through the large unsaturated zone of the CT aquifers substantially reduces the tritium concentration reaching the saturated zone (Brennwald et al., 2001). \(^4\)He\(^3\)He is different for each sample and can be used as a qualitative groundwater age as in Figure 3. If the PR-model is applied instead of the CE-model one obtains \((\HeHe)_{\text{atm}} = (3.0 \pm 0.1) \times 10^{-8}\) and \(\HeHe = 11 \pm 1\) TU. Both values are significantly larger than those predicted by the CE-model. Including the \(^20\)Ne\(^{22}\)Ne ratios in the data set, 70 observables are available for the 42 fit parameters leaving 28 degrees of freedom. For this data set the predicted values of \(\HeHe\) and \(\HeHe\) are the same as in case of the PR-model. Nevertheless the calculated values of \(\HeHe\) and \(\HeHe\) are used as best estimates.

Figure 6a displays the predicted concentration of non-atmospheric \(^4\)He, i.e., the sum of \(\HeHe\) and \(\HeHe\), versus the predicted \(\HeHe\) for the PR- and the CE-model obtained from inverse fitting of \(T, A, F\) or \(T, A, F\) and \(F\), respectively, to the elemental concentrations of \(\Ne, \Ar, \Kr\), and \(\Xe\). The regression lines give in case of the CE-model \((\HeHe)_{\text{atm}} = (2.2 \pm 0.1) \times 10^{-8}\) and \(\HeHe = 0.4 \pm 0.2\) TU and in case of the PR-model \((\HeHe)_{\text{atm}} = (3.4 \pm 0.1) \times 10^{-8}\) and \(\HeHe = 12 \pm 1\) TU. Fitting regression lines corresponds to the second approach taken above and confirms the results provided by the ensemble fitting using Noble.

In previous studies (e.g., Aeschbach-Hertig et al., 2000; Kipfer et al., 1994; Kipfer et al., 1996) \((\HeHe)_{\text{atm}}\) has been derived from a regression line in a three isotope plot \(^3\)He/\(^4\)He vs Ne/\(^4\)He (Fig. 6b). Such a procedure assumes that all measured ratios can be interpreted as a mixture between a terrigene component with constant \((\HeHe)_{\text{ter}}\) and \(\NeHe = 0\), and a combined atmospheric/tritiogenic component with constant composition given by \((\HeHe)_{\text{atm}}\) and \(\NeHe\). Although these assumptions are not strictly fulfilled, the variations in the atmospheric/tritiogenic endmember (due to varying \(T, A, F\), and \(\HeHe\)) are relatively small compared to the effect of the terrigene component. Therefore this approach yields a useful estimate of \((\HeHe)_{\text{ter}}\).
4.2.3. Estimation of water age of young groundwater samples

The estimated values of \(( ^4\text{He}_{\text{ter}}/^{4}\text{He}_{\text{atm}})\), \(^3\text{He}_{\text{atm}}\) and \(^4\text{He}_{\text{atm}}\) from the different models can be employed to estimate \(^3\text{He}_{\text{tr}}\), in young groundwater samples containing tritium. The resulting \(^3\text{He}_{\text{tr}}\) differs between models and so does the \(^3\text{H}-^3\text{He}\) water age \(\tau\) (Tolstikhin and Kamenskiy, 1969; Torgersen et al., 1979):

\[
\tau = \frac{1}{\lambda} \ln \left( 1 + \frac{^3\text{He}_{\text{tr}}}{^3\text{H}} \right)
\]

where \(\lambda = 0.05626\) yr\(^{-1}\) (Lucas and Unterweger, 2000). Depending on the choice of the gas exchange model, \(^3\text{H}-^3\text{He}\) water ages differ by up to a factor of 2 (Fig. 5b). Shown are the water ages for all samples in which \(^3\text{H}\) has been detected and \(^3\text{He}_{\text{tr}}\) is significantly larger than zero. The water ages predicted by the PR-model appear unrealistically large. They suggest that around 1930 the tritium concentration at recharge \((^3\text{H}_{\text{rech}} = ^3\text{H}_m + ^3\text{He}_{\text{tr}})\), i.e., at the bottom of the 30 m thick unsaturated zone, were up to 10 TU, which seems unrealistic. A consistency test of the \(^3\text{H}-^3\text{He}\) water ages based on the comparison of \(^3\text{H}_{\text{rech}}\) with the historic \(^3\text{H}\) concentration at recharge (Aeschbach-Hertig et al., 1998; Stute et al., 1997) is not practicable in the CT-aquifers because the \(^3\text{H}\) at recharge is strongly affected by the transport through the unsaturated zone. Hence, only the inclusion of the Ne isotopes in the analysis provides an objective, i.e., data-based, criterion to identify the \(^3\text{H}-^3\text{He}\) water ages predicted by the CE-model to be more reliable than those predicted by the PR-model.

4.2.4. Simplified procedures to estimate \(^4\text{He}_{\text{atm}}\)

In many studies the dating procedure outlined above is not applicable because measurements of heavy noble gases are not available. Ne can be used to estimate excess air even if the recharge temperature is only approximately known, because the equilibrium concentration \(^1\text{Ne}\) only weakly depends on temperature (Table 1). Usually (e.g., Aeschbach-Hertig et al., 1998; Ekwurzel et al., 1994; Schlosser et al., 1989; Torgersen et al., 1979) the Ne excess \((\Delta \text{Ne} = \text{Ne}_{\text{atm}} - \text{Ne})\) has been used to calculate the atmospheric He excess from \(\Delta \text{He} = \Delta \text{Ne} \cdot (z_{\text{He}}/z_{\text{Ne}})\), i.e., assuming unfractionated excess air (UA-model). However, if the excess air is fractionated, its \(^1\text{He}/\text{Ne}\) ratio can be significantly lower than \(z_{\text{He}}/z_{\text{Ne}}\) (Holocher et al., 2001), particularly in the case of the PR-model.

The equivalent application of the UA-model to the Ne data from young groundwater samples \((\text{Ne}_{\text{atm}} < 1.5 \times 10^{-7} \text{ cm}^3\text{STPg}^{-1})\) of the CT-aquifers, assuming the recharge temperature \(T\) to correspond to the annual mean soil temperature of \(\sim 32\,^\circ\text{C}\) (CESBIO/ORSTOM/CNES.HAPEX SAHEL Information System, 1996), results in predicted concentrations of \(^4\text{He}_{\text{atm}}\) that are 0.7 to 3.2% larger than those predicted by the CE-model using the concentrations of all noble gases and the \(^{20}\text{Ne}/^{22}\text{Ne}\) ratio (Fig. 5a, Table 3). This demonstrates that the estimation of \(^4\text{He}_{\text{atm}}\) based on \(\Delta \text{Ne}\) and the assumption of unfractionated excess air provides a reasonable approximation if the gas excess is not strongly fractionated and fractionation is not of PR-model type. However, the information contained in the Ne concentration is not sufficient to test the validity of the assumption that fractionation is small or to distinguish between the different models.
It is possible to estimate the effect of fractionation on $^4\text{He}_{\text{atm}}$ and $^3\text{He}_{\text{atm}}$ if the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio is considered in addition to the elemental Ne concentration. Using these two constraints and assuming that recharge temperature is approximately known, the model Eqn. 1 and 2 can be solved for the parameters $A_{pr}$ and $F_{pr}$ or $A_{ce}$ and $F_{ce}$, respectively. With the derived model parameters, $^4\text{He}_{\text{atm}}$ and $^3\text{He}_{\text{atm}}$ can be calculated from Eqn. 1 and 2, respectively. Equivalently, Noble can be employed to fit the parameters $A_{pr}$ and $F_{pr}$ or $A_{ce}$ and $F_{ce}$, respectively, based on the measured Ne concentration, the $^{20}\text{Ne}/^{22}\text{Ne}$ ratios and a prescribed recharge temperature.

Application of this method to the Ne concentrations and $^{20}\text{Ne}/^{22}\text{Ne}$ ratios from the CT aquifers provides $^3\text{He}_{\text{atm}}$ and $^4\text{He}_{\text{atm}}$, which are very similar for all three models (PR-, CE- and UA-model). The reason is that the inverse fitting based only on Ne concentrations and the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio suggests that fractionation is small. The deviation between the $^3\text{He}_{\text{atm}}$ predicted using the reduced data set and the best estimate of $^3\text{He}_{\text{atm}}$ predicted using all available data ranges between $-3.9\%$ and $3.2\%$ for all young groundwater samples with $^4\text{He}_{\text{tot}} < 1.5 \times 10^{-7}$ cm$^3$STPg$^{-1}$ (see Table 3). Although a clear decision between UA-, CE- and PR-model is impossible on the basis of Ne concentration and $^{20}\text{Ne}/^{22}\text{Ne}$ ratio, the range of possible $^3\text{He}_{\text{atm}}$ can be estimated and strong fractionation of PR-model type can be distinguished from fractionation according to the CE-model or from unfractonated excess air as in the UA-model (Table 3).

Compared to the PR-model, the He-Ne fractionation according to the CE-model is not very different from unfractonated excess air (UA-model). Hence for dating studies the distinction between the CE- and the UA-model is not as critical as the distinction between the CE-and the PR-model. Nevertheless, in the case of samples with very young age, for which the difference between total $^3\text{He}$ and $^3\text{He}_{\text{atm}}$ is small, and in the case of a large contribution of excess gas, the prediction of $^3\text{He}_{\text{atm}}$ and water age may differ significantly depending on whether CE- or UA-model is applied. Since the UA-model corresponds to the CE-model with $F_{ce} = 0$, the distinction between CE- and UA-model essentially is the problem of determining the parameters $A_{ce}$ and $F_{ce}$ of the CE-model from a limited data set on noble gases and noble gas isotopes. The $^{20}\text{Ne}/^{22}\text{Ne}$ ratio is not suitable to determine whether $F_{ce} = 0$ or not, because the solubilities of the Ne isotopes are very similar such that the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio in atmospheric equilibrium is approximately the same as in the atmosphere. However, estimation of $A_{ce}$ and $F_{ce}$ could be based on the Ar concentration together with the Ne concentration and the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio. The Ne/Ar ratio in atmospheric equilibrium differs significantly from that in the atmosphere (Table 2) and Ar concentrations can be measured comparatively easy and are less dependent on recharge temperature than the concentrations of the heavy noble gases Kr and Xe.

In case of the Niger data, the constraint by Ne and Ar concentrations together with the $^{20}\text{Ne}/^{22}\text{Ne}$ ratios is sufficient to rule out the UA- and PR-models, whereas the predictions according to the CE-model are compatible with the data from the young groundwater samples according to a $\chi^2$-test ($p(X^2_{\text{test}} = 29) > 3\%$). Considering the groundwater samples individually, predictions according to the CE-model are consistent with the data of all samples ($p(X^2) > 1\%$ for all 27 samples) while the predictions of the PR- and the UA-model are acceptable only for 2 and 3 samples, respectively. Note that including the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio is required to distinguish between the PR- and CE-model. Fitting based on Ne and Ar concentrations alone only allows rejection of the UA-model (Table 3).

The parameters $A_{ce}$ and $F_{ce}$ estimated from inverse fitting based on Ne and Ar concentrations and the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio can be employed to determine the concentration of $^3\text{He}_{\text{atm}}$ and $^4\text{He}_{\text{atm}}$. The $^3\text{He}_{\text{atm}}$ concentrations agree within $\pm 0.2\%$ with the reference value obtained from the concentrations of Ne, Ar, Kr, Xe and the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio (Table 3). In Figure 5a the predicted $^3\text{He}_{\text{atm}}$ (crosses) is compared to the predictions using the UA-model based only on Ne (triangles) and to the predictions using the CE- and the PR- models based on concentrations of Ne, Ar, Kr and Xe (solid and open circles, respectively).

The prediction of $^3\text{He}_{\text{atm}}$ based on the concentrations of Ne, Ar and the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio only weakly depends on $T$. For recharge temperatures between 30°C and 35°C $^3\text{He}_{\text{atm}}$ deviates from the best value by less than 1% for the young groundwater samples (Table 3). The $\chi^2_{\text{ref}}$ for the fits with $T = 30°C$ and $T = 35°C$ are larger than for $T = 32°C$, indicating that the latter is the most appropriate value for the recharge temperature, in agreement with the the soil temperature and the noble gas temperatures derived from all noble gas concentrations using the CE-model.

$^3\text{H}-^3\text{He}$ water ages calculated from the $^3\text{He}_{\text{atm}}$ and $^4\text{He}_{\text{atm}}$ obtained from the simplified procedures discussed above agree well with the $^3\text{H}-^3\text{He}$ water ages derived from the entire data set using the CE-model (Fig. 5b). Water ages derived using the usual He correction based only on $\Delta$Ne are typically 1 to 2 yr younger than the best estimates. This is particularly important in studies concerned with dating in very young groundwaters. In the samples 19 and 30 $^3\text{H}$ and $^4\text{He}_{\text{atm}}$ are very small, such that small uncertainties in $^3\text{He}_{\text{atm}}$ lead to large errors in the $^3\text{H}-^3\text{He}$ water age. Water ages based on Ne, Ar and $^{20}\text{Ne}/^{22}\text{Ne}$ agree with the best estimates within error and in most cases the $^3\text{H}-^3\text{He}$ water ages deviate by less than 0.2 yr.

5. CONCLUSIONS

Concentrations of noble gas isotopes carry information on gas partitioning during groundwater recharge which might not be available from the elemental noble gas concentrations of Ne, Ar, Kr, and Xe. Specifically, in the CT-aquifers of Niger a decision between the PR- and the CE-model is impossible based only on the elemental concentrations. If however measurements of the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio are included in the analysis, the PR-model has to be rejected according to a $\chi^2$-test and only the CE-model is compatible with the entire data set. This finding is also confirmed by the $^{36}\text{Ar}^{40}\text{Ar}$ ratio, which is incompatible with the PR-model. Thus the concentrations of noble gas isotopes allow a decision between PR- and CE-model which influences the prediction of recharge temperature, initial amount of excess air, and fractionation and hence affects the interpretation of the data with respect to paleoclimatic conditions.

The distinction between PR-, CE- and UA-model is also important in $^3\text{H}-^3\text{He}$ dating. In groundwaters with fractionated
gas excess, $^3\text{He}_{\text{atm}}$ required to calculate $^4\text{He}_{\text{atm}}$ strongly depends on the fractionation model employed. Commonly the excess gas contributing to $^4\text{He}_{\text{atm}}$ is estimated from $\Delta\text{Ne}$ assuming unfractonated excess air (UA-model). The effect of fractionation can only be assessed by including additional experimental information. Combination of the Ne concentration with the $^{20}\text{Ne}/^{22}\text{Ne}$ ratio allows estimation of the range of possible values of $^4\text{He}_{\text{atm}}$ thus providing the range of possible water ages. However, this procedure is not suitable to distinguish between PR-, CE- and UA-model if fractionation is small, i.e., if the $^{20}\text{Ne}/^{22}\text{Ne}$ is between the value in the atmosphere and the value at atmospheric equilibrium. Extending the data set consisting of Ne concentration and $^{20}\text{Ne}/^{22}\text{Ne}$ ratio by the concentration of Ar enables a rigorous distinction between models and a reliable estimation of $^4\text{He}_{\text{atm}}$ and $^4\text{He}_{\text{atm}}$ of $^3\text{H}-^3\text{He}$ water ages.

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